And schedule data releases

Felipe Almeida-Fernandes

### **Summary**

## **Calibration Pipeline**

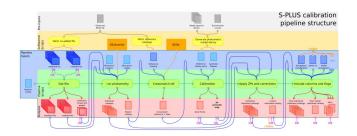
Detection
Apertures
Aperture Correction
Reference Catalogs
Model Fitting Calibration
Stellar Locus Calibration
Results

#### **Data Releases**

Schedule Catalogs Format VACs

## Felipe **Almeida-Fernandes** et al. 14th S-PLUS meeting - December 2020

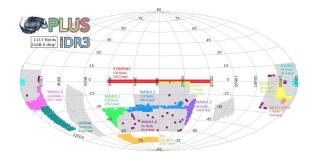
## **Calibration**



## The S-PLUS calibration pipeline

- Technique developed for S-PLUS
- New features include:
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  - ISM extinction included in template fitting
  - Gaia scale calibration
  - CCD ZP offsets correction
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## **Data Releases**



## Internal data release 3 - iDR3

- Already available for members of the collaboration
- 1818.8 sq deg (+ 289.7)
- >50 Million detections

## Public Data Release 2 - DR2

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- March 2021
- 950.5 sq deg
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## **Summary**

## **Calibration Pipeline**

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## The Calibration Pipeline

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## **Summary**

### **Calibration Pipeline**

Detection

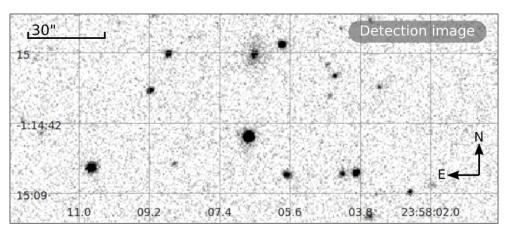
Apertures
Aperture Correction
Reference Catalogs
Model Fitting Calibration
Stellar Locus Calibration
Results

### **Data Releases**

Schedule Catalogs Format VACs

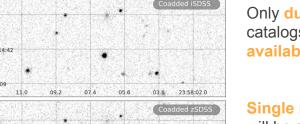
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## **Detection Image**



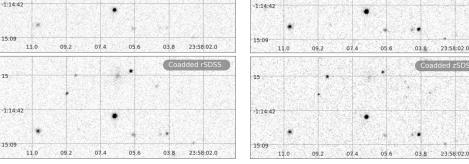
Detection image obtained from filters: G, R, I, Z

**SExtractor** is used both for source **detection** and **photometry**.



Only dual mode catalogs are currently available.





## **Summary**

#### **Calibration Pipeline**

Detection

#### **Apertures**

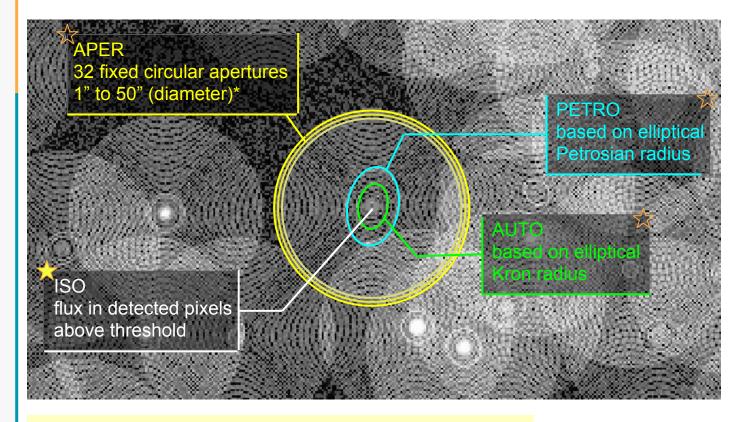
Aperture Correction
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#### **Data Releases**

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## **Photometry - Different Apertures**



\* Of the fixed apertures, only 3" and 6" are given in the final catalogs

## **Summary**

### **Calibration Pipeline**

Detection

Aperture Correcti

## Aperture Correction

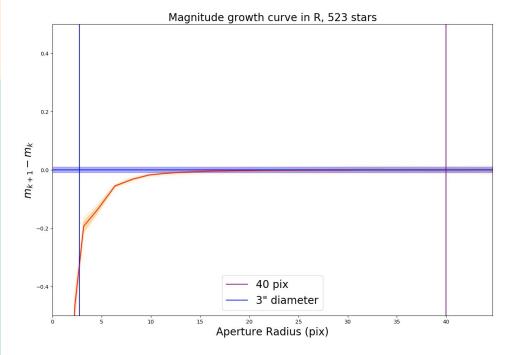
Reference Catalogs Model Fitting Calibration Stellar Locus Calibration

Results

### **Data Releases**

Schedule Catalogs Format VACs

## **★** Aperture Correction for Point Sources



For any given aperture, a fraction of the flux is lost due to the PSF

Aperture corrections are obtained for each field and filter

Hundreds of stars are used in each case



### **Summary**

## **Calibration Pipeline**

Detection
Apertures
Aperture Correction

## Reference Catalogs

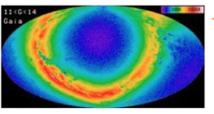
Model Fitting Calibration Stellar Locus Calibration Results

#### **Data Releases**

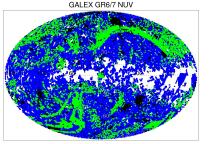
Schedule Catalogs Format VACs

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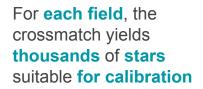
## **Reference Catalogs**



ATLAS Refcat2
All-Sky
Tonry et al. 2018



GALEX GR6/GR7
Almost All-Sky
Bianchi et al. 2017



 An advantage of the large FoV of the S-PLUS images



**SDSS Standard Stars Catalog for STRIPE82\*** 

Only STRIPE82 lvezić et al. 2007

\*12 STRIPE 82 fields are not covered by Ivezic's catalog. SDSS DR14 is used instead

## **Summary**

### **Calibration Pipeline**

Detection
Apertures
Aperture Correction
Reference Catalogs

**Model Fitting Calibration** 

Stellar Locus Calibration

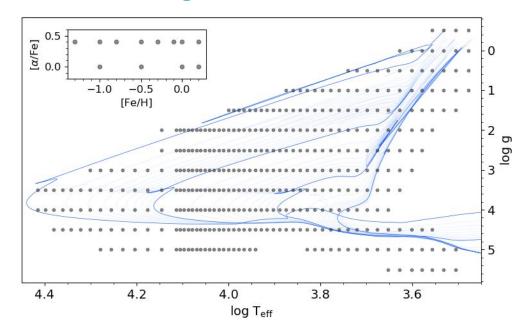
Results

### **Data Releases**

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## **Model Fitting - Theoretical Models** ★



Theoretical Stellar Templates of Coelho 2014:



Grid that covers the whole parameters space



ISM extinction included as a free parameter



Pre-convoluted to the filter systems of:

SDSS, PanSTARRS, GAIA, GALEX, Skymapper, DES and S-PLUS



## **Summary**

### **Calibration Pipeline**

Detection Apertures Aperture Correction Reference Catalogs

**Model Fitting Calibration** 

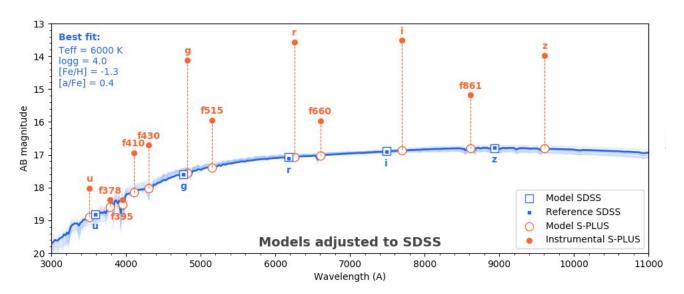
Stellar Locus Calibration Results

#### **Data Releases**

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## **Model Fitting - External Calibration**



 Best template for each star is derived by minimizing the chi2
 First, using the reference magnitudes.

$$\chi^2_{\text{ext},i} = \sum_{m \in \mathcal{R}} \left( \frac{m_{\text{mod},i} - m_{\text{ref}}}{\delta m_{\text{ref}}} \right)^2$$

 Templates used to predict S-PLUS calibrated magnitudes

## **Summary**

### **Calibration Pipeline**

Detection Apertures Aperture Correction Reference Catalogs

**Model Fitting Calibration** 

Stellar Locus Calibration

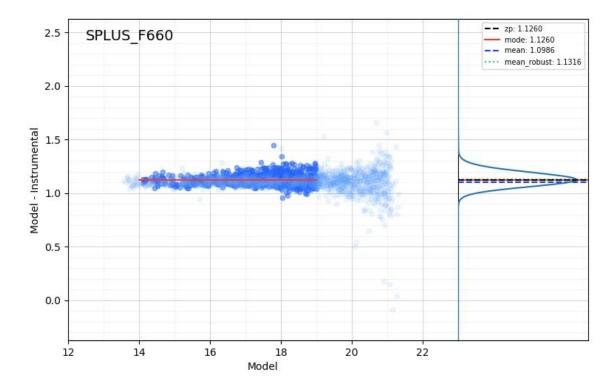
Results

### **Data Releases**

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## **Model Fitting - External Calibration**



- Photometric Zero Points are derived from the difference between instrumental and model predicted S-PLUS magnitudes.
  - Obs: an instrumental ZP of 20mag is already included

## **Summary**

### **Calibration Pipeline**

Detection Apertures Aperture Correction Reference Catalogs

## Model Fitting Calibration

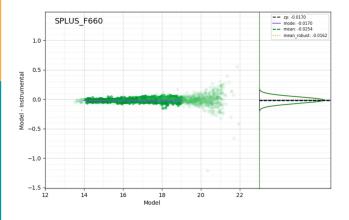
Stellar Locus Calibration Results

#### **Data Releases**

Schedule Catalogs Format VACs

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## **Model Fitting - Internal and Gaia Calibration**

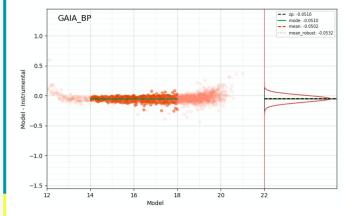


## **Internal Calibration**



We fit models to the S-PLUS calibrated magnitudes to refine the calibration

 Use of 12 bands to better constrain the templates



## **Calibration to Gaia Standard**



We fit models to the S-PLUS calibrated magnitudes to predict Gaia magnitudes

 The Gaia offsets are added to bring S-PLUS calibration to the same magnitude scale of Gaia

## **Summary**

### **Calibration Pipeline**

Detection **Apertures Aperture Correction** Reference Catalogs Model Fitting Calibration

#### Stellar Locus Calibration

Results

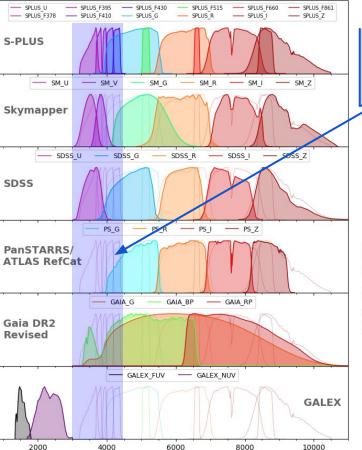
#### **Data Releases**

Schedule **Catalogs Format VACs** 

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## Model Fitting - Stellar Locus Calibration ★

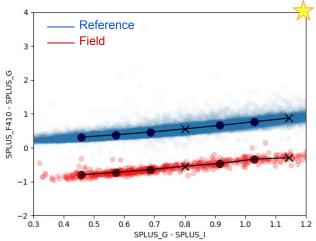




Wavelength (A)

If ATLAS RefCat is the only reference available for the field, the blue bands are not properly covered.

5 Bluer bands are calibrated using an S-PLUS reference Stellar Locus:





## **Summary**

### **Calibration Pipeline**

Detection
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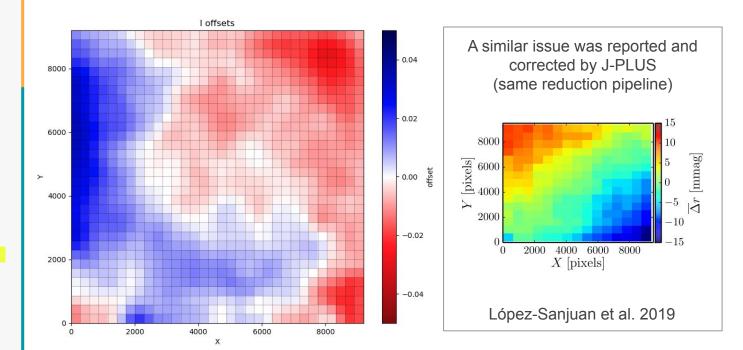
### **Data Releases**

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## CCD correlated corrections \*



We find that fluctuations around the ZP correlates with the source's XY position in the CCD



Correction maps are produced and applied as the first step after the photometry.



## **Summary**

### **Calibration Pipeline**

Detection
Apertures
Aperture Correction
Reference Catalogs
Model Fitting Calibration
Stellar Locus Calibration

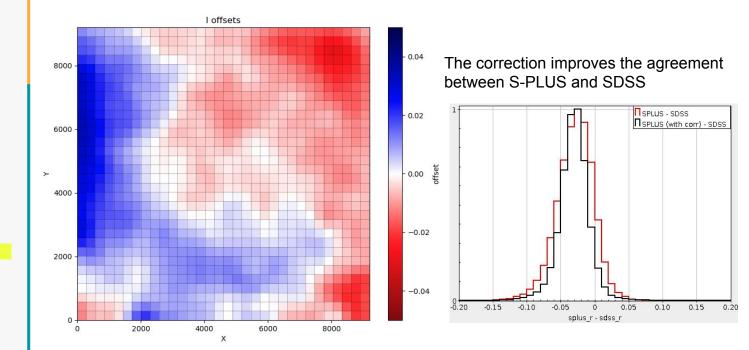
### **Data Releases**

Results

Schedule Catalogs Format VACs

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## **Summary**

## **Calibration Pipeline**

Detection

**Apertures** 

**Aperture Correction** 

Reference Catalogs

Model Fitting Calibration

Stellar Locus Calibration

Results

## **Data Releases**

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## Results

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## **Summary**

## **Calibration Pipeline**

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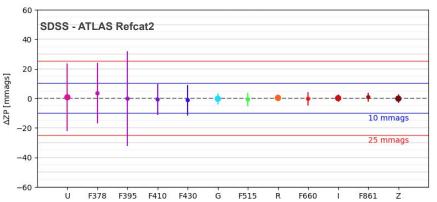
#### **Data Releases**

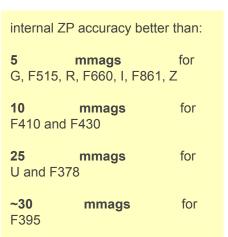
Schedule Catalogs Format VACs

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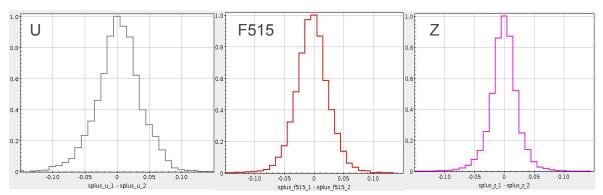
## **Internal Consistency**

## ZP comparison using different reference catalogs





## Comparison between overlapping fields



## **Summary**

### **Calibration Pipeline**

Detection
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Stellar Locus Calibration

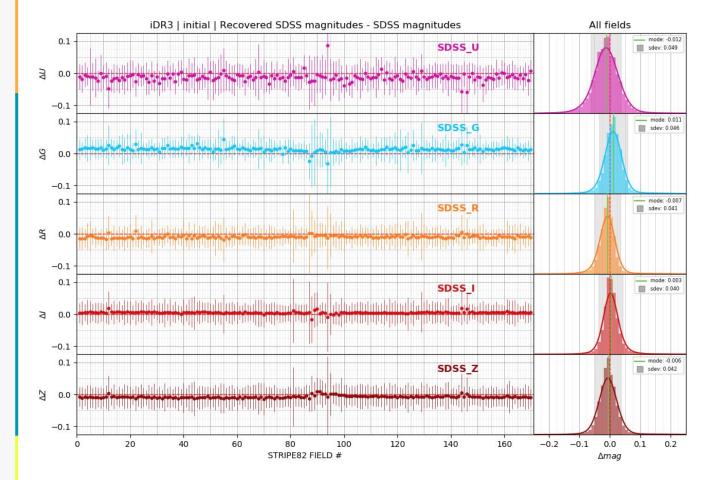
Results

### **Data Releases**

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## **External Consistency - SDSS**



## **Summary**

Detection

### **Calibration Pipeline**

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### **Data Releases**

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## **Data Releases**

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## **Summary**

Catalogs Format **VACs** 

**Calibration Pipeline** Detection **Apertures** Aperture Correction Reference Catalogs Model Fitting Calibration Stellar Locus Calibration Results **Data Releases** Schedule

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Released

Scheduled

## DR1 iDR2 iDR3

DR2

DR3

iDR4

iD

## Dec/19 Sep/20

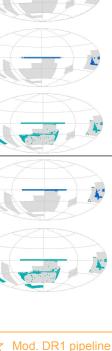
Late 21

Release

**Date** 

**Type** 

Internal



Current pipeline

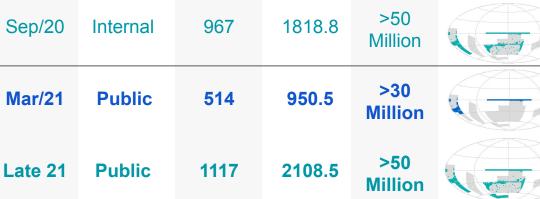
RA, DEC



**TBD** 

No

**Fields** 



**TBD** 

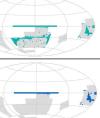
**Million** ~25 Million

[Sq deg]

All S-PLUS Data Releases No Area **Footprint** 

**Detections** 

**TBD** 



## **Summary**

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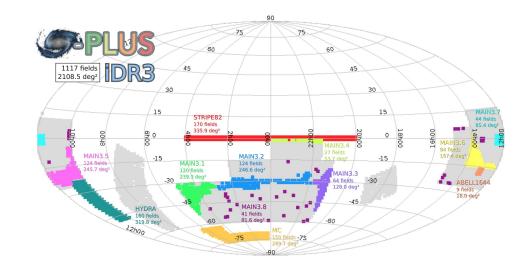
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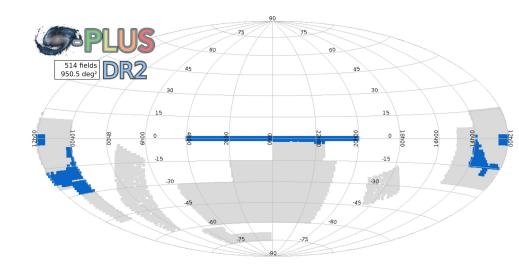
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**VACs** 

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## DR2 and iDR3 catalogs format

The final product of the pipeline are the calibrated multi-filter photometric catalogs of each field.

Each catalog has 283 columns, including astrometry, PSF morphology and photometry in six different apperture with respective errors and S/N.

IDs and position	Detection parameters	
Field ID RA DEC	X Y ISOarea MU_MAX A	KRON_RADIUS PhotoFlagDet CLASS_STAR FWHM s2n_Det_auto s2n_Det_petro
Calibration flag	THETA ELONGATION	s2n_Det_jetto s2n_Det_iso s2n Det aper 3
calibration_flag	ELLIPTICITY FLUX_RADIUS	s2n_Det_aper_6 FWHM_n

## Individual filters photometry and parameters

```
PhotoFlag_{filter}
For each filter:
                      CLASS_STAR_{filter}
U, F378, F395,
                      FWHM {filter}
F410, F430,
                      {filter}_{aper}
                                            For each aperture:
G, F515,
                                            AUTO, PETRO, ISO,
                      e {filter} {aper}
R, F660, I
                                            APER 3, APER 6
                      s2n_{filter}_{aper}
F861, Z
                                            PSTOTAL
                      nDet_{aper}
```

## **Summary**

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VACs

## **VACs**

The information in the pipeline's output catalogs is basically **astrometry + photometry** 

The value added catalogs will include columns that require a subsequent analysis of the data. VACs that will be available soon are:

## STAR/GALAXY/QUASAR Classification

Lilianne Nakazono

## **Photometric Redshifts**

Erik Vinicius

## **Data quality flags**

Maria Luiza Buzzo

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### **Summary**

## **Calibration Pipeline**

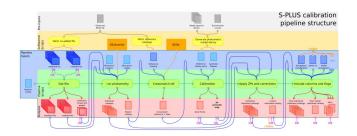
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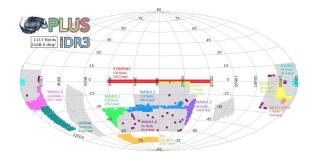
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