

**Data Release 2 of S-PLUS** 

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15th S-PLUS meeting - 01-03 Jun 2021

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les covering an area of

the new generation of

to already calibrated

%. We compare the

we show that these ZPs

tter constrain the models by adding

rs, which are ≤ 10 mmags for filters J0410,

 $x_{r}$ , and  $\leq 25$  mmags for filters u and J0395. We describe

#### Data Release 2 of S-PLUS: accurate template-fitting based photometry covering ~1000 square degrees in 12 optical filters

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#### ABSTRACT

The Southern Photometric Local Universe Survey (S-PLUS) is an ongoing survey of ~9300 deg2 photometric system. This paper presents the second data release (DR2) of S-PLUS, submitted to MNRAS 950 deg2. The data has been fully calibrated using a new photometric calibrated wide-field multi-filter surveys. This technique consists of a  $\chi^2$  minim data from other surveys, eliminating the need for stand template-predicted and S-PLUS instrumental can be further refined by fitting the the narrow-band information J0430, g. J0515, r. J0660 ns to the final catalogues and present a brief characterisation of the the complete data flow of t data. We show that, for a r rse threshold of 3, the photometric depths of the DR2 range from 19.9 mag to 21.3 mag (measured in F rtures), depending on the filter. The S-PLUS DR2 can be accessed from the website: https://splus.cloud.

Key words: surveys - techniques: photometric - catalogues - astronomical data bases: miscellaneous - stars: general - galaxies: general

#### 1 INTRODUCTION

Wide-field photometric surveys are essential for research in astronomy, especially because of the large volume of data they are able to provide in a reasonable amount of time and with more extensive sky coverage compared to spectroscopic surveys. Surveys such as the Sloan Digital Sky Survey (SDSS, York et al. 2000), 2MASS (Skrutskie et al. 2006). ATLAS (Shanks et al. 2015), and PanSTARRS, (Chambers et al. 2016), to mention a few, have contributed to the development of countless areas in astronomy: from the study of asteroids to the large scale structure of the Universe.

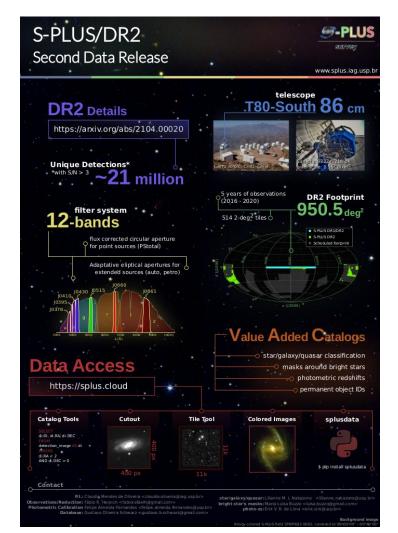
Following the success of these past surveys, several ongoing and planned projects are being executed to complement and supplement

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the available data in terms of increasing i) the sky-coverage; mainly by including the southern hemisphere (e.g. DES, Abbott et al. 2018; Skymapper Wolf et al. 2018); ii) the photometric-depth: reaching fainter magnitudes (e.g. LSST; lvezić et al. 2019); or iii) the wavelength range and resolution: extending or increasing the number of pass-bands in previous filter systems and even replacing broadband with narrow-band filters to widen the spectral feature sensitivity (e.g. the Pristine Survey, Starkenburg et al. 2017).

Regarding the topic of expanding the wavelength resolution, three surveys clearly stand out: the Javalambre Physics of the Accelerating Universe Astrophysical Survey (JPAS, Benitez et al. 2014; and miniJ-PAS, Bonoli et al. 2020), the Javalambre Photometric Local Universe Survey (J-PLUS, Cenarro et al. 2019) and the Southern Photometric Local Universe Survey (S-PLUS, Mendes de Oliveira et al. 2019). Of these, J-PAS is the most ambitious and plans to cover an area



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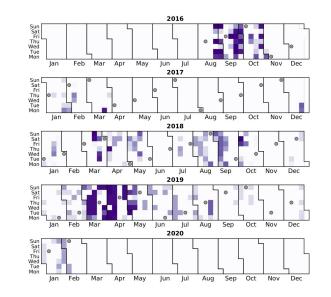
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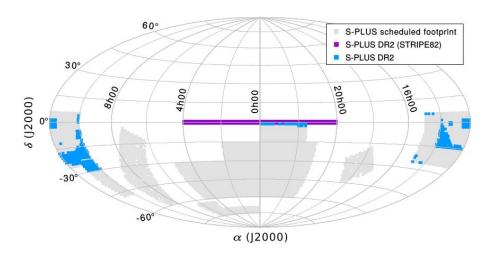
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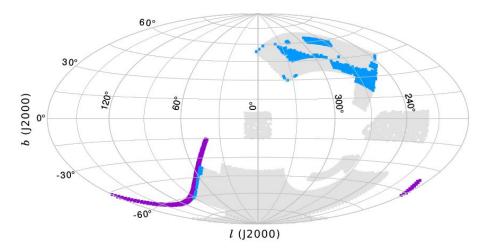
**9-PLUS** 

### • 514 S-PLUS fields

- 950.5 sq degrees
- Observations between 2016 and 2020







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Data

Ide	entification and Position		
Field	Name of the S-PLUS field		
	of the observation		
ID	Observation ID in DR2		
RA	Right Ascension (J2000)	[deg] [deg]	
DEC	Declination (J2000)		
X	CCD X-axis position		
	(reduced image)	[pixel]	
Y	CCD Y-axis position		
	(reduced image)	[pixel]	
	Morphology		
<b>ISOarea</b>	Number of pixels exceeding		
	1.1 sigma threshold		
MU_MAX	Peak surface brightness	[mag/arcsec2]	
	above background		
A	Isophotal image	[pixel]	
	major axis		
В	Isophotal image	[pixel]	
	minor axis		
THETA	Isophotal image	[deg]	
	position angle		
ELONGATION	A/B		
ELLIPTICITY	1 - B/A		
FLUX_RADIUS	Radius containing	[pixel]	
	(0.2,0.5,0.7,0.9)		
	fraction of the light		
KRON_RADIUS	Kron apertures		
	in units of A or B		
FWHM	FWHM assuming a	[pixel]	
	Gaussian core		
FWHM_n	Normalized FWHM		
FWHM {filter}	FWHM at each filter	[pixel]	

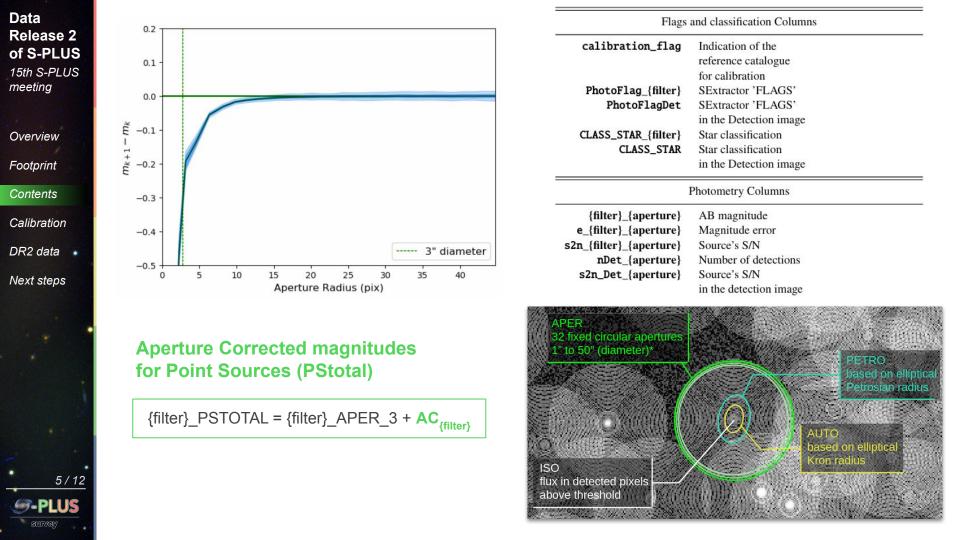
reference catalogue for calibration SExtractor 'FLAGS' SExtractor 'FLAGS' in the Detection image
SExtractor 'FLAGS' SExtractor 'FLAGS'
SExtractor 'FLAGS'
in the Detection image
Star classification
Star classification
in the Detection image
Photometry Columns
AB magnitude
Magnitude error
Source's S/N
Number of detections
Source's S/N
Source s S/IN

- photometric redshifts
- masks around bright stars
- permanent S-PLUS ID

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SULLYON
activesy

Field	Name of the S-PLUS field	
	of the observation	
ID	Observation ID in DR2	
RA	Right Ascension (J2000)	[deg]
DEC	Declination (J2000)	[deg]
х	CCD X-axis position	
	(reduced image)	[pixel]
Y	CCD Y-axis position	
	(reduced image)	[pixel]
	Morphology	
IS0area	Number of pixels exceeding	
	1.1 sigma threshold	
MU_MAX	Peak surface brightness	[mag/arcsec <sup>2</sup> ]
	above background	
A	Isophotal image	[pixel]
	major axis	
В	Isophotal image	[pixel]
	minor axis	
THETA	Isophotal image	[deg]
	position angle	
ELONGATION	A/B	
ELLIPTICITY	1 - B/A	
FLUX_RADIUS	Radius containing	[pixel]
	(0.2,0.5,0.7,0.9)	
	fraction of the light	
KRON_RADIUS	Kron apertures	
	in units of A or B	
FWHM	FWHM assuming a	[pixel]
	Gaussian core	
FWHM_n	Normalized FWHM	

	Flags a	and classification Columns
	calibration_flag	Indication of the reference catalogue
	PhotoFlag_{filter} PhotoFlagDet CLASS_STAR_{filter}	for calibration SExtractor 'FLAGS' SExtractor 'FLAGS' in the Detection image Star classification
	CLASS_STAR	Star classification in the Detection image
	I	Photometry Columns
	{filter}_{aperture} e_{filter}_{aperture} s2n_{filter}_{aperture} nDet_{aperture} s2n_Det_{aperture}	AB magnitude Magnitude error Source's S/N Number of detections Source's S/N in the detection image
	APER 32 fixed circular apertures 1" to 50" (diameter)*	PETRO based on elliptical Petrosian radius
ISC		AUTO based on elliptical Kron radius
abo	k in detected pixels	



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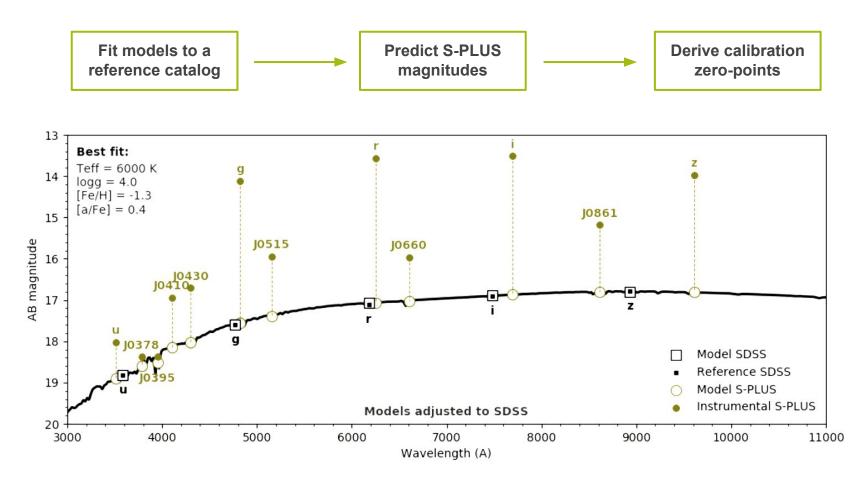
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## **Calibration approach:**



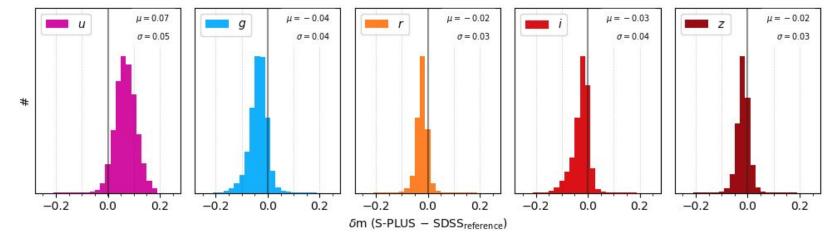
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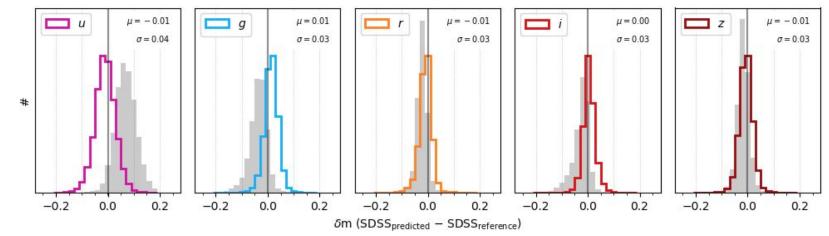
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#### Comparison after converting to SDSS filter system



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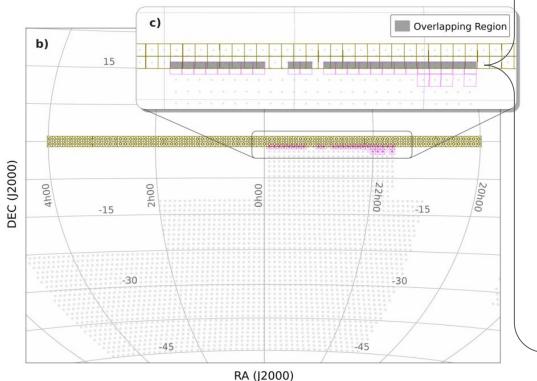
Calibration DR2 data

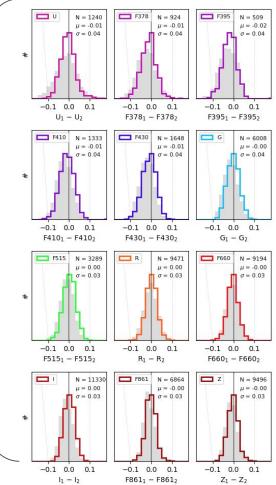


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# Internal comparison of overlapping fields





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Characterization of DR2 data

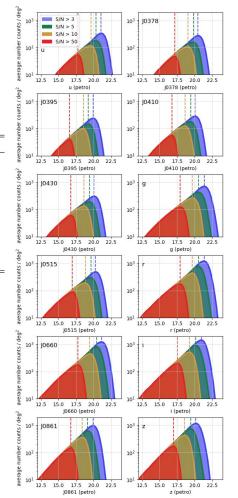
 Photometric depths

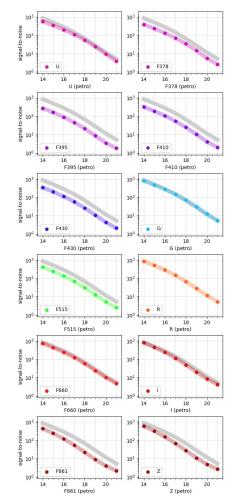
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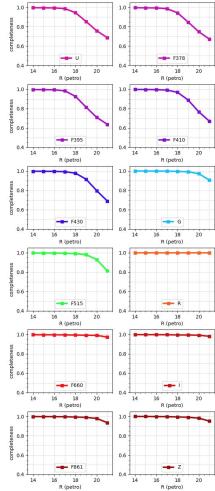
fi

ilter	S/N > 50	S/N > 10	S/N > 5	S/N > 3
ı	17.7	19.6	20.3	21.0
0378	17.1	19.0	19.8	20.4
0395	16.5	18.5	19.2	19.9
0410	16.7	18.6	19.4	20.0
0430	16.7	18.6	19.4	20.0
	17.9	19.6	20.5	21.3
0515	16.9	18.8	19.6	20.2
	17.9	19.6	20.5	21.3
0660	17.7	19.4	20.4	21.1
	17.5	19.2	20.1	20.9
0861	16.7	18.3	19.1	19.9
	16.9	18.5	19.4	20.1

- Average s/n per magnitude bin
- Completeness in relation to the r-band







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## S-PLUS is suitable for many different sciences

It is very challenging to provide quality data for so many different applications will remain the highest resolution large photometric survey in this area in the near future.

The extended resolution of these surveys, covering key spectral features, enables their application in many different fields. In particular, S-PLUS data has been used to study clusters of galaxies, considering accurate photometric redshifts in Stripe-82 using template fitting (Molino et al. 2020) and machine learning (Vinicius-Lima et al., submitted), ultra-diffuse galaxies (Barbosa et al. 2020), lenticular galaxies (Cortesi et al. 2021), the Hydra cluster galaxies (Lima-Dias et al. 2021), conduct searches for guasars (Nakazono et al. submitted), determine galaxy morphology (Bom et al. submitted), perform star/galaxy separation (Costa-Duarte et al. 2019), analyse stellar populations in and around the Milky Way, including stellar groups in the CMa OB1 association (Santos-Silva et al. submitted), determine and study the photometric metallicity and carbon distributions of stars in the Milky Way's Halo (Whitten et al. accepted), find and characterize compact planetary nebulae (Gutiérrez-Soto et al. 2020) and ultra metal-poor stars (Placco et al., in prep) and study active low-mass stars in CMa R1 star-forming region (Gregorio-Hetem et al. 2021), as well as several ongoing projects. Given this wide range of applications, it is of utmost importance to provide precise and accurate photometry that is reliable for both point and extended sources.

The photometric calibration is the process of translating photon counts measured at the detector into physical fluxes at the top of the

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Photometry

Inclusion of single mode photometry

Support to weight images

Inclusion of psf photometry (DOPHOT)

New detection images generated by

#### **Aperture Correction**

**Recent improvements** 

Detection

SWARP

Refined parameters for moderately crowded fields

#### Calibration (work in progress)

- Improvements in the model fitting process
- Extinction correction for MC and low galactic latitudes

#### Data Access

- new database: splus.cloud
- new python package: splusdata
- support to Topcat TAP service

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#### **Recent improvements**

Detection

- New detection images generated by SWARP
- Support to weight images
- Photometry
  - Inclusion of single mode photometry
  - Inclusion of psf photometry (DOPHOT)

- Aperture Correction

 Refined parameters for moderately crowded fields



